

# WCSTools 3.0

## Tools for Image Astrometry and Catalog Searching

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### Astronomical Catalogs Supported by WCSTools

Catalog or Format	No. Stars	Bytes	Region Search	Image Search
USNO-A2.0 Catalog	526,280,881	6,315,370,572	scat (scat)	imms2 (imcat)
GSC II Catalog	455,851,237	7	gspec (scat)	imgsc2 (imcat)
2MASS Point Source Catalog	155,509,249	48,988,432,908	stmc (scat)	imsmc (imcat)
GSC-A2.0 Catalog	55,368,239	664,418,868	asm2 (scat)	imsm2 (imcat)
USNO-AGT Catalog	25,541,952	1,231,787,520	agsca (scat)	imgscA (imcat)
HST Guide Star Catalog	25,541,952	1,231,787,520	agsca (scat)	imgsc (imcat)
Tycho-2 Catalog	2,539,913	528,721,576	sty2 (scat)	imty2 (imcat)
USNO/Hipparcos ACT Catalog	988,758	318,380,076	scat (scat)	imcat (imcat)
PPM Catalog	378,910	22,734,656	sppm (scat)	imppm (imcat)
SAO Catalog	258,996	15,539,876	sas (scat)	immsa (imcat)
IRAS Point Source Catalog	245,889	7,376,698	siras (scat)	imiras (imcat)
Hipparcos Catalog	118,218	449,232	shp (scat)	imshp (imcat)
Starbase tab-delimited ASCII	varies	varies	scat	imcat
Space-Delimited ASCII	varies	varies	scat	imcat
Local Binary	varies	varies	scat	imcat
USNO-A1.0 Catalog	488,006,860	5,856,082,320	uaal (scat)	imual (imcat)
USNO-S1.0 Catalog	54,787,624	657,451,488	su1 (scat)	imsu1 (imcat)
USNO-J1.0 Catalog	19,911,514	238,938,168	sujc (scat)	imsujc (imcat)

### The GSC II Catalog over the World Wide Web

All WCSTools programs which access the GSC II Catalog use the web. Code to assemble and send http queries to the STScI GSC II server is built into the GSC II access subroutine. Since tab-separated Starbase tables, a native WCSTools catalog format, are an output option from the catalog server, they are parsed and sorted just like local Starbase files.

The following example searches a 3 arcminute radius circle around 10 00 30.00 J2000 and returns all the stars found there, sorted by their distance from the search center. The verbose mode (-v) echoes all the http query text.

```
$ wget -e -s -H 180 10 00 30 00 J2000
wget 3 0 5 21 September 2001, Doug Mink SAO
Center: 10 00 00.000 +30 00 00.00 J2000
Area: 00 00 00 27 31 x 00 00 00 00
http://tdc-www.harvard.edu/cgi-bin/gsc2/gsc2?ra=10+00+00.00&dec=30+00+00.00&radius=180.00&sort=dist&format=starbase&verbose=1
http://tdc-www.harvard.edu/cgi-bin/gsc2/gsc2?ra=10+00+00.00&dec=30+00+00.00&radius=180.00&sort=dist&format=starbase&verbose=1
#starbase tab-delimited ASCII
#USNO-A2.0 Catalog
#GSC II Catalog
#2MASS Point Source Catalog
#GSC-A2.0 Catalog
#USNO-AGT Catalog
#Tycho-2 Catalog
#USNO/Hipparcos ACT Catalog
#PPM Catalog
#SAO Catalog
#IRAS Point Source Catalog
#Hipparcos Catalog
#Starbase tab-delimited ASCII
#Space-Delimited ASCII
#Local Binary
#USNO-A1.0 Catalog
#USNO-S1.0 Catalog
#USNO-J1.0 Catalog
```

Running the same search with a vvv flag prints the actual returned information, (the returned lines are split here to fit on the page).

```
$ wget -vvvr 180 10 00 30 00 J2000
#starbase tab-delimited ASCII
#USNO-A2.0 Catalog
#GSC II Catalog
#2MASS Point Source Catalog
#GSC-A2.0 Catalog
#USNO-AGT Catalog
#Tycho-2 Catalog
#USNO/Hipparcos ACT Catalog
#PPM Catalog
#SAO Catalog
#IRAS Point Source Catalog
#Hipparcos Catalog
#Starbase tab-delimited ASCII
#Space-Delimited ASCII
#Local Binary
#USNO-A1.0 Catalog
#USNO-S1.0 Catalog
#USNO-J1.0 Catalog
```

### SCAT used over the World Wide Web

scat can call itself over the web. Code to turn http query input into the same parameters which are set on the command line is built into the main scat program. Code to assemble and send http queries to remote servers, such as the ESO USNO-A2.0 and GSC servers and the STScI GSC II server, as well as remote scat servers, is built into the individual catalog access subroutines. Since tab-separated Starbase tables, a native WCSTools catalog format, are an output option on most catalog servers, they are parsed and sorted just like local Starbase files.

```
$ wget -vvvr 180 10 00 30 00 J2000
#starbase tab-delimited ASCII
#USNO-A2.0 Catalog
#GSC II Catalog
#2MASS Point Source Catalog
#GSC-A2.0 Catalog
#USNO-AGT Catalog
#Tycho-2 Catalog
#USNO/Hipparcos ACT Catalog
#PPM Catalog
#SAO Catalog
#IRAS Point Source Catalog
#Hipparcos Catalog
#Starbase tab-delimited ASCII
#Space-Delimited ASCII
#Local Binary
#USNO-A1.0 Catalog
#USNO-S1.0 Catalog
#USNO-J1.0 Catalog
```

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#USNO-A2.0 Catalog
#GSC II Catalog
#2MASS Point Source Catalog
#GSC-A2.0 Catalog
#USNO-AGT Catalog
#Tycho-2 Catalog
#USNO/Hipparcos ACT Catalog
#PPM Catalog
#SAO Catalog
#IRAS Point Source Catalog
#Hipparcos Catalog
#Starbase tab-delimited ASCII
#Space-Delimited ASCII
#Local Binary
#USNO-A1.0 Catalog
#USNO-S1.0 Catalog
#USNO-J1.0 Catalog
```

### WCSTools Documentation:

<http://tdc-www.harvard.edu/software/wcstools/>

This poster in HTML:

<http://tdc-www.harvard.edu/software/wcstools/publications/adass2001>

This poster in Postscript:

<http://tdc-www.harvard.edu/software/wcstools/publications/adass2001/poster.ps>

### An Evolving Software Package

For five years, WCSTools as provided image astrometry for astronomers who need accurate positions for objects they wish to observe. Other functions have been added and improved since the package was first released.

► **imwvs** has been improved in several ways:

- > Fits may be repeated, with and without tightening the tolerance on matches between image and catalog stars.
- > Any of several magnitudes in a reference catalog can be used to order the brightest stars for matching.
- > A WCS can be fit to an initial set of X-Y/RA-Dec matches, with optional full match and fit of additional image stars.

► **WCSTools** includes and uses Mark Calabretta's latest release of **WCSTools** version 2.7, for projections, and supports multiple world coordinate systems in a single header.

► **Catalogs** are now searchable over the World Wide Web. The catalog searching routine can be located on either end (or both ends!) of such a web connection, and the output from one catalog search can be used as the input to another search.

► **Support** has been added for new catalogs, such as the GSC-ACT and 2MASS Point Source Catalog, as they have come out. Access over the web to the GSC II Catalog Server at the Space Telescope Science Institute is supported as a default. A simple command line interface can search any supported catalog, returning information in several standard formats.

► **scat** has been made to work more like a filter, so the output of a search of one catalog can be used as input for a search of another catalog. It has also been extended so it can be used as part of a web site which serves archived spectra.

The goal of this project is to have a good set of command line tools for dealing with images, their headers, source catalogs, and relationships between them. The FITS WCS standard will be completely supported when it is agreed upon, and access to more catalog-serving web sites will be added in the future.

### Multiple World Coordinate Systems

WCSTools implements the use of multiple world coordinate systems, with the WCS to which a keyword applies designated by a suffixed upper-case alphabetic character.

This **CTYPE1** applies to the default WCS, which may be named using **WCSENAME**. A second WCS could use keywords such as **CTYPE1A** and **WCSENAMEA**. Up to 27 (no suffix, A-Z) different sets of WCS keywords may be specified.

In the WCSTools programs, the WCS to be used is specified by adding **WCSENAME** to the end of a filename specified on the command line. This is added after the extension of a multi-extension FITS image is specified.

`fitsfile.extension_name_or_number:WCSENAME_or_letter`

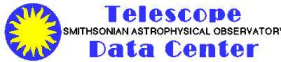
**imwvs** always sets the default (no suffix) keywords.

Sometimes getting from image pixels to world coordinates is best thought of as a two step process, `pixel->focal plane->sky`, for example. The keyword **WCSENAME**, indicating a named WCS conversion which must be applied to pixel coordinates before WCS conversion x can be applied, could be added to indicate the name or letter of the requisite WCS.

### Example of Multiple WCS in a FITS Header

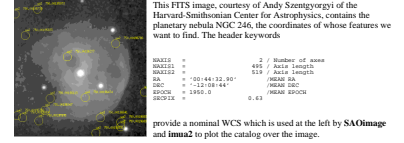
Thanks to Steve Allen, Lick Observatory

```
CRVAL1 = 0.5
CRVAL2 = 89.5
CRPIX1 = 1
CRPIX2 = 1
CRD1_1 = 1
CRD1_2 = -1
WCSENAME = 'TLOW-CAR' / Terrestrial East Longitude
CTYPE1 = 'TLOW-CAR' / Terrestrial North Latitude
CUNIT1 = 'deg'
CUNIT2 = 'deg'
WCSENAME = 'image' / redundant restoration of default FITS
CRPIX1 = 1 / reference pixel along FITS axis 1
CRPIX2 = 1 / reference pixel along FITS axis 1
CRVAL1 = 1 / WCS value along axis 1 at refpix
CRVAL2 = 1 / WCS value along axis 2 at refpix
CD1_1 = 1 / CTM 1,1 from pixel to WCS
CD1_2 = 0 / CTM 1,2 from pixel to WCS
CD2_1 = 0 / CTM 2,1 from pixel to WCS
CD2_2 = 1 / CTM 2,2 from pixel to WCS
CTYPE1 = 'IMEX-CAR' / coordinate/projection type of axis 1
CTYPE2 = 'IMGY-CAR' / coordinate/projection type of axis 2
CUNIT1 = 'pixel' / physical unit along WCS axis 1
CUNIT2 = 'pixel' / physical unit along WCS axis 2
```



### IMWCS Using Prematched stars and the USNO-A2.0 Catalog

If an image has features which make it hard to match specific stars, it is useful to match up a few image stars (5 is a reasonable number) to catalogued positions by hand. **imwvs**, as of version 2.8.7, accepts a file of X,Y,RA,Dec matched positions and fits an initial world coordinate system to them. This WCS may be either written into the image header or used as an initial guess to fit a more complete list of image stars positions and a deep catalog, such as the US Naval Observatory A2.0 Catalog, to get a better WCS.



Because the nebulosity confuses automatic star-finding programs and the reference catalogs, a prematched set of x, y, right ascension, and declination coordinates for 5 stars in the field were placed in the file `ngc246.stars`:

```
72.04 143.88 00 47 10.884 -11 53 18.92
114.03 185.80 00 47 09.153 -11 52 52.70
370.15 400.11 00 46 58.055 -11 50 36.60
325.49 181.03 00 47 00.064 -11 52 49.89
418.23 30.94 00 46 56.044 -11 54 32.13
```

These selected stars are plotted in the image at the right. These matches are used by **imwvs** to fit a WCS using header and command-line information for an initial value.

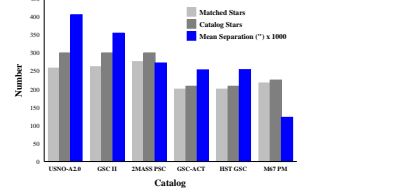
To quickly find the 100 brightest stars in the image, **Emmanuel Bernin's Extractor** program was used through a shell script, **imx** which sets up the parameter files and prepares the image, if necessary:

```
$ imms2 -new -d ngc246.star -u ngc246.stars ngc246.fits
$ imms2 -new -d ngc246.star -u ngc246.stars ngc246.fits
#starbase tab-delimited ASCII
#USNO-A2.0 Catalog
#GSC II Catalog
#2MASS Point Source Catalog
#GSC-A2.0 Catalog
#USNO-AGT Catalog
#Tycho-2 Catalog
#USNO/Hipparcos ACT Catalog
#PPM Catalog
#SAO Catalog
#IRAS Point Source Catalog
#Hipparcos Catalog
#Starbase tab-delimited ASCII
#Space-Delimited ASCII
#Local Binary
#USNO-A1.0 Catalog
#USNO-S1.0 Catalog
#USNO-J1.0 Catalog
```

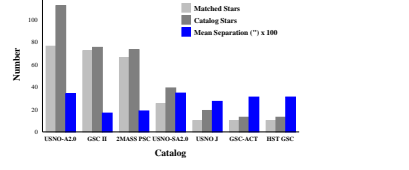
The final FITS image contains standard FITS world coordinate system keywords and a logging keyword showing what version of **imwvs** it was:

```
WCSNAME = 'IMEX-CAR' / coordinate/projection type of axis 1
WCSNAME = 'IMGY-CAR' / coordinate/projection type of axis 2
WCSNAME = 'IMEX-CAR' / coordinate/projection type of axis 1
WCSNAME = 'IMGY-CAR' / coordinate/projection type of axis 2
WCSNAME = 'IMEX-CAR' / coordinate/projection type of axis 1
WCSNAME = 'IMGY-CAR' / coordinate/projection type of axis 2
WCSNAME = 'IMEX-CAR' / coordinate/projection type of axis 1
WCSNAME = 'IMGY-CAR' / coordinate/projection type of axis 2
WCSNAME = 'IMEX-CAR' / coordinate/projection type of axis 1
WCSNAME = 'IMGY-CAR' / coordinate/projection type of axis 2
```

### IMWCS Fit to M67 Cluster Field



### IMWCS Fit to Random Field



### SCAT in the Virtual Observatory

scat can be used as a filter in a CGI script as well as running as a CGI program itself. In this case, it helps search and serve a catalog of spectra and redshifts from the Smithsonian's Z-Machine reticon spectrograph which was run on our 60 inch telescope on Mt. Hopkins in Arizona from 1978 through 1993. Output from scat is run through an awk script which adds the links to CGI scripts using the ID number found by scat. The links call a Fortran program to plot a spectrum, a C program to list information about the spectrum, and IRAF SPP programs, through CL shell scripts, to return FITS or ASCII pixel list files. The entire process is managed by a Perl script.

